



Introduction

Relevance

- Capture reaction ${}^{14}C(n, \gamma){}^{15}C$ controls the carbon-nitrogen-oxygen cycle in the helium-burning regions of stars [1]
- Coulomb dissociation [1] and direct capture [2] data are compared with theory predictions
- Effectiveness of multiple EFT expansions are tested

RESEARCH OBJECTIVES

- The ${}^{14}C(n, \gamma){}^{15}C$ cross section is calculated, and a systematic expansion is obtained [3]
- Ambiguity regarding parameters sizes leads to multiple theoretical expressions
- Bayesian analysis is applied to 1) determined unknown parameters and 2) compare the evidences for each theoretical expression

Cross Section

The ¹⁴C(n, γ)¹⁵C reaction rate is reflected in the cross section:

$$\sigma(p) = \frac{1}{2} \frac{64\pi\alpha}{M_c^2 \mu^2} \frac{p\gamma(p^2 + \gamma^2)}{1 - \rho\gamma} \left[2 \left| g^{2P_{1/2}}(p) \right|^2 + 4 \right| g^{2P_{1/2}}(p) \right]^2 + 4 \left| g^{2P_{1/2}}(p) \right|^2 + 4 \left| g^{2P_{1/2}}(p) \right$$

where

$$g^{2}P_{3/2}(p) = \frac{\mu}{p^2 + \gamma^2} + \frac{6\pi\mu}{1/a_1^{(2)} + r_1^{(2)}p^2/2 - s_1^{(2)}p^4/4 - ip^3} \left[\frac{\gamma}{4\pi} + \frac{6\pi\mu}{1/a_1^{(2)} + r_1^{(2)}p^2/2 - s_1^{(2)}p^4/4 - ip^3}\right]$$

unknown parameters in red [3]

- The cross section is expressed as an expansion in Q/Λ :
- Low-momentum scale Q~40 MeV
- High-momentum scale ∧~100-200 MeV
- Two different models derived based on different sets of assumptions (TABLE 1)
- LO and NLO expansions for each model

 TABLE 1. Parameter sizes

Model A Model B $\rho \sim \frac{1}{Q}$ $\rho \sim \frac{1}{\Lambda}$ $a_1^{(2)} \sim \frac{1}{\Lambda Q^2}$ $a_1^{(2)} \sim \frac{1}{Q^3}$ $r_1^{(2)} \sim \Lambda$ $r_1^{(2)} \sim Q^3$ $s_1^{(2)} \sim \frac{1}{\Lambda}$

 $6\pi(p^2+\gamma^2)$

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Bayesian Methods for ¹⁴C Neutron Capture

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- [8] R Reifarth et al., Journal of Physics **41**, 053101 (2014)



[7] K. Barbary, https://github.com/kbarbary/nestle